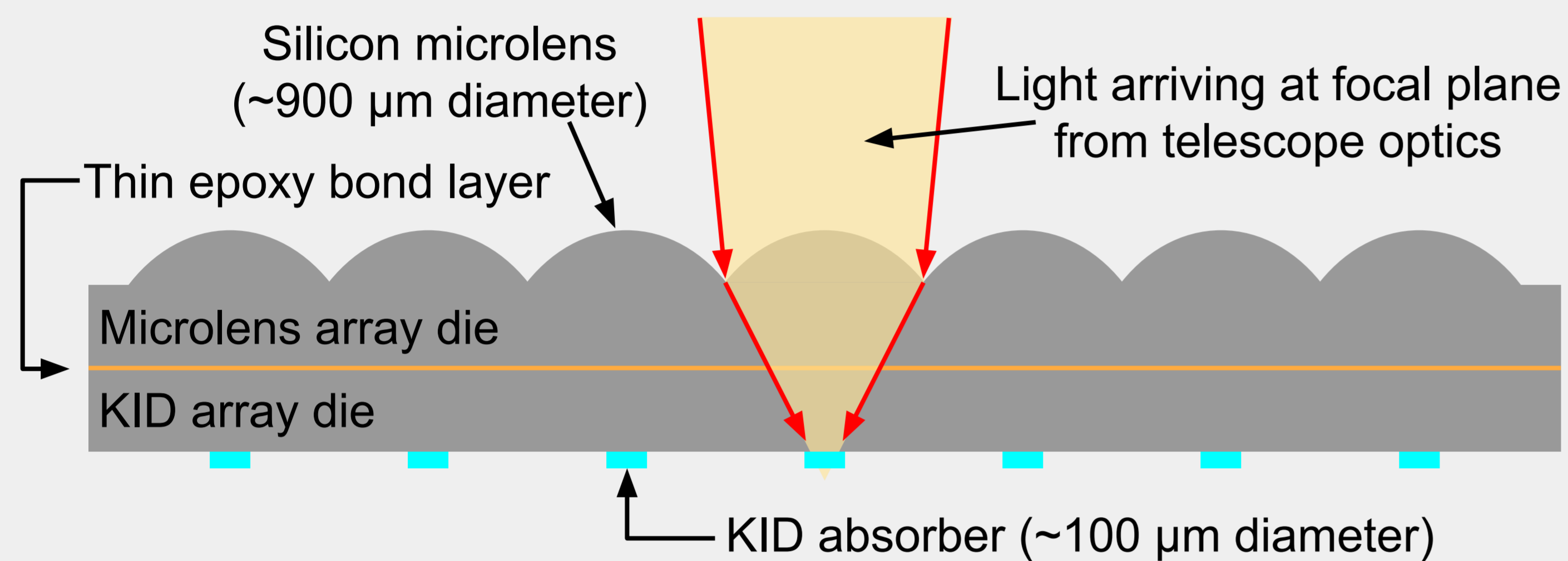


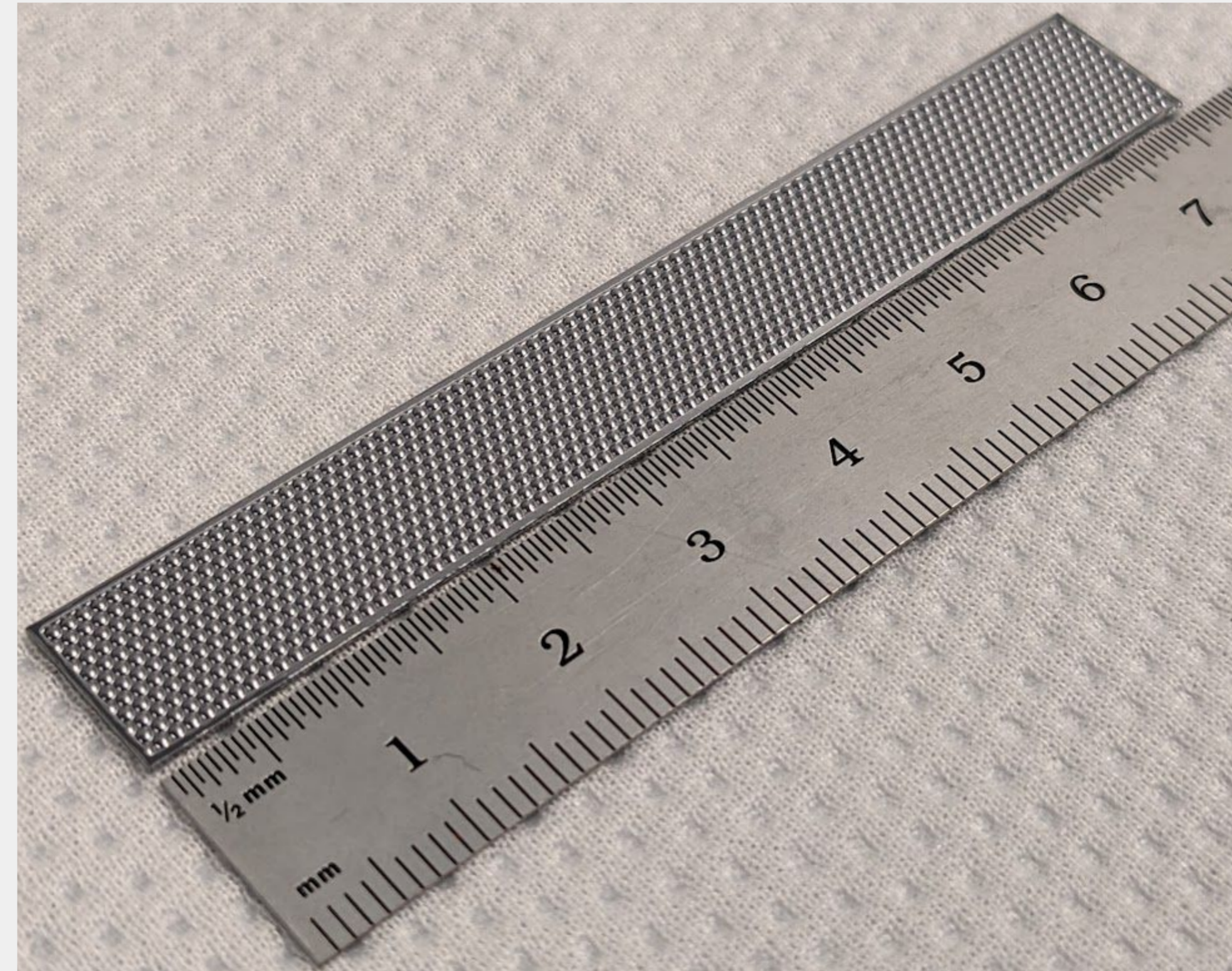
## Far-IR Monolithic Silicon Microlens Arrays

- The PRobe far-Infrared Mission for Astrophysics (PRIMA)
  - Observations of AGN feedback, interstellar medium physics in galaxies
  - Growth over cosmic time of stars, black holes, heavy elements, dust
  - Far-IR Kinetic Inductance Detector (KID) arrays [1,2,3]
  - Many arrays, each with thousands of detectors,  $\lambda \sim 24 - 240 \mu\text{m}$
- Silicon monolithic microlens arrays for KID arrays
  - One microlens for each pixel
  - Concentrate optical power on to detector
- Fabrication methods developed:
  - Lithography-based microfabrication
  - Detector-microlens array hybridization

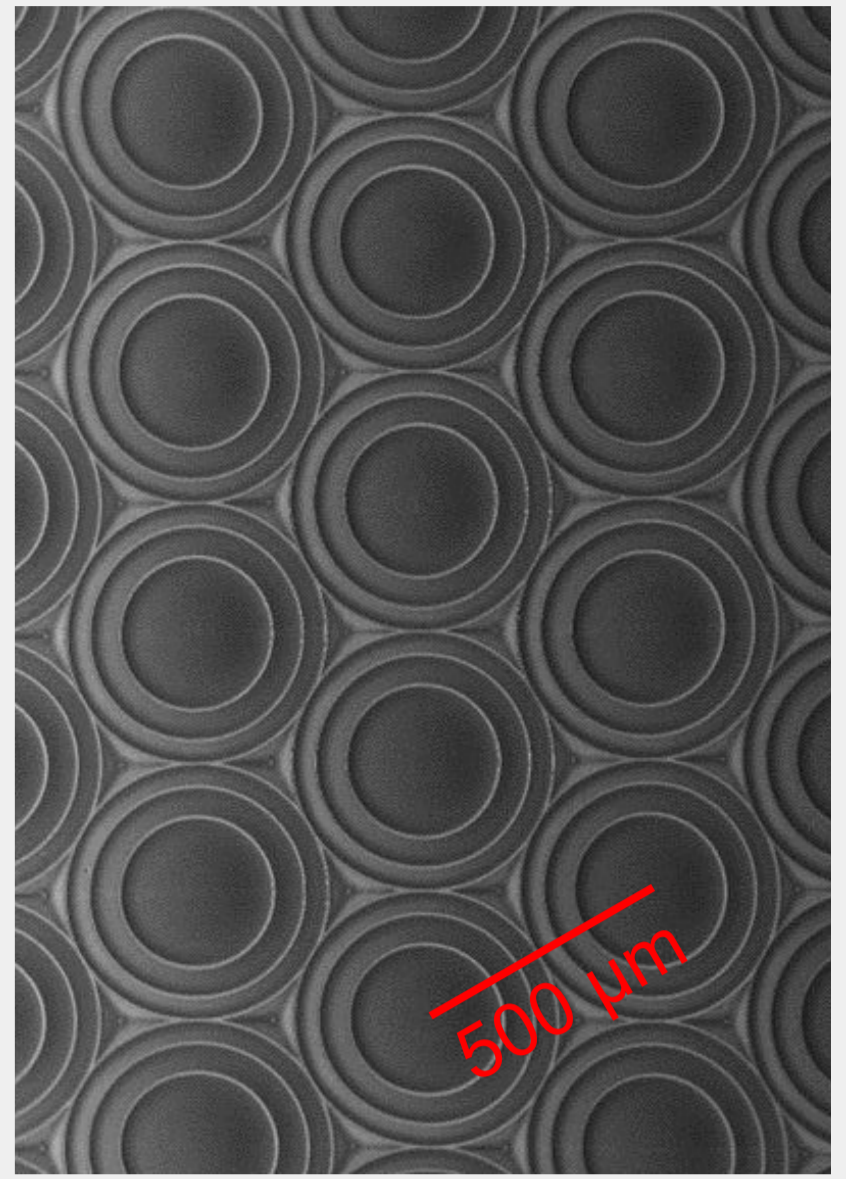


## Fabrication of Microlens Arrays

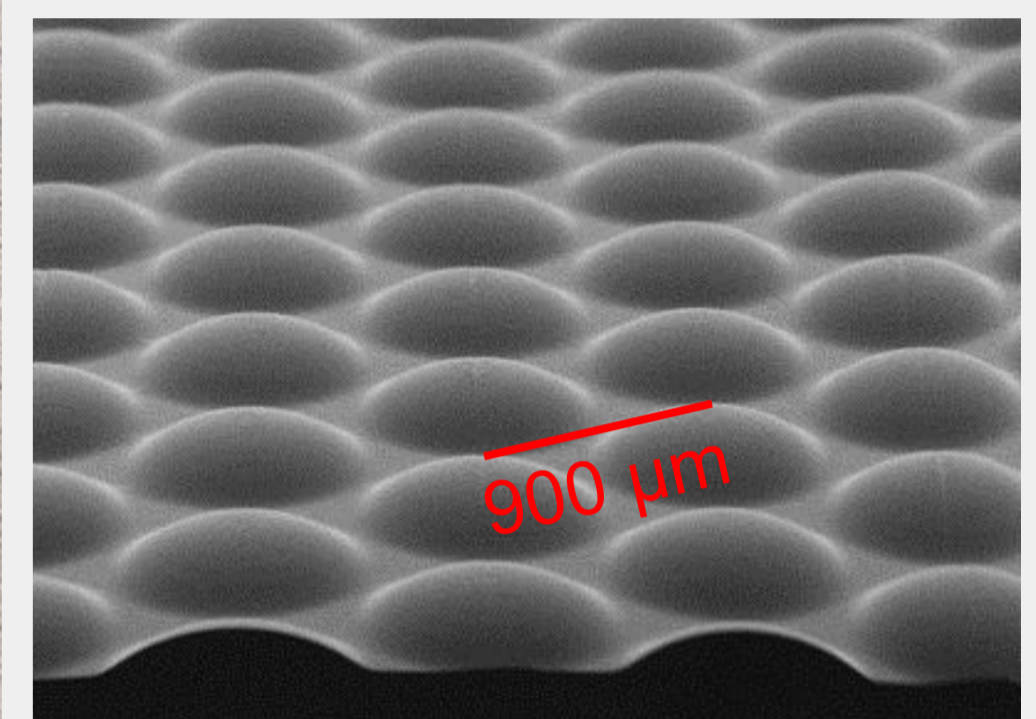
- Microlens array patterned with grayscale lithography
  - Calibrate dosage via exposure-contrast curve
- Plasma etching to transfer lens profile into silicon
  - Control etch selectivity to achieve desired profile
- Successful demonstration of lens depths of  $\sim 150 \mu\text{m}$ 
  - Microlens pitches in the range  $500 - 900 \mu\text{m}$
  - Fresnel-style and full-sag lenses



Left: Kilo-pixel full-sag microlens array for  $\lambda \sim 200 \mu\text{m}$ .



SEM image of  $\lambda \sim 25 \mu\text{m}$  Fresnel microlens array.



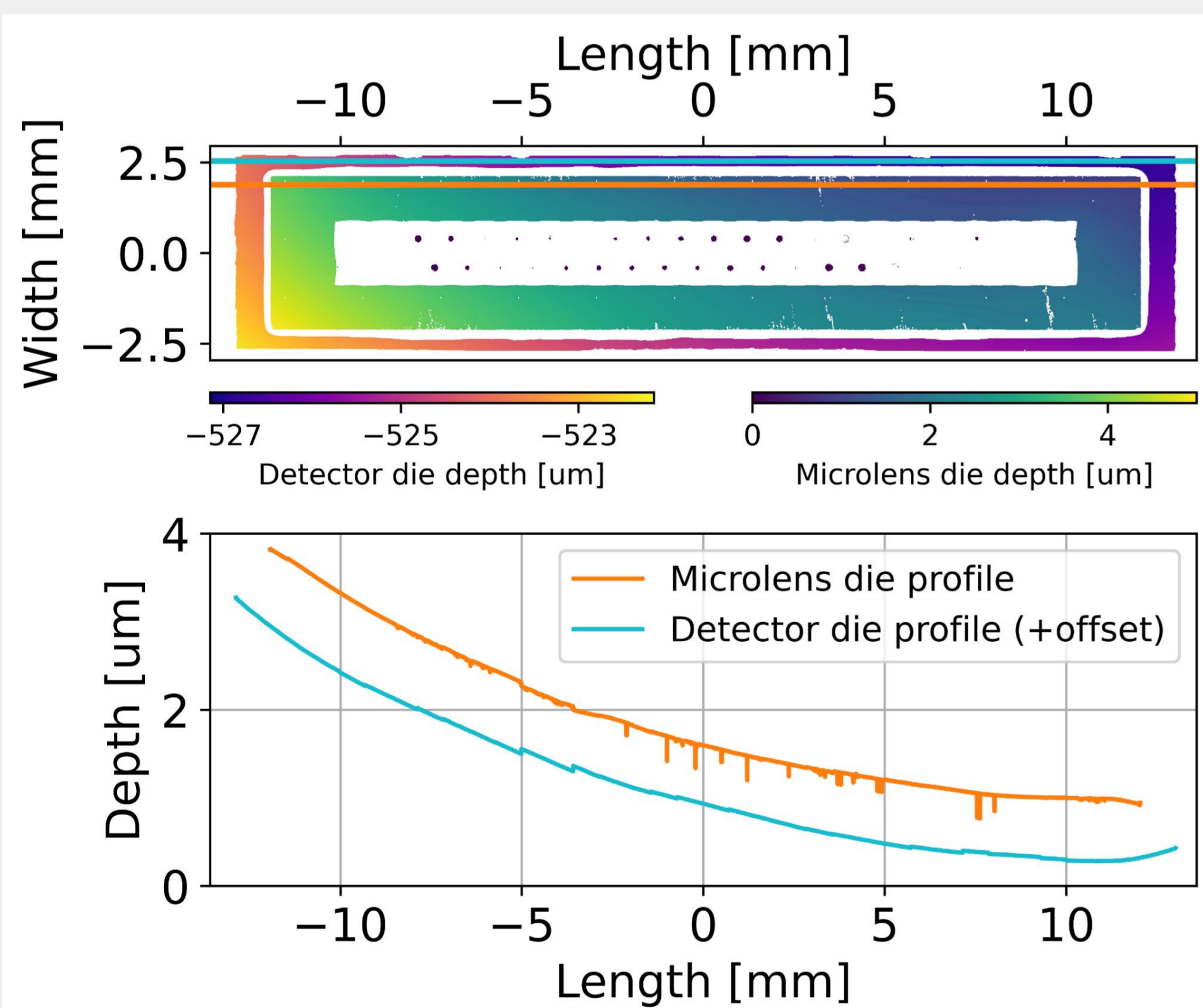
SEM image of  $\lambda \sim 200 \mu\text{m}$  full-sag microlens array.

## Hybridization of Microlens Arrays with Detector Arrays

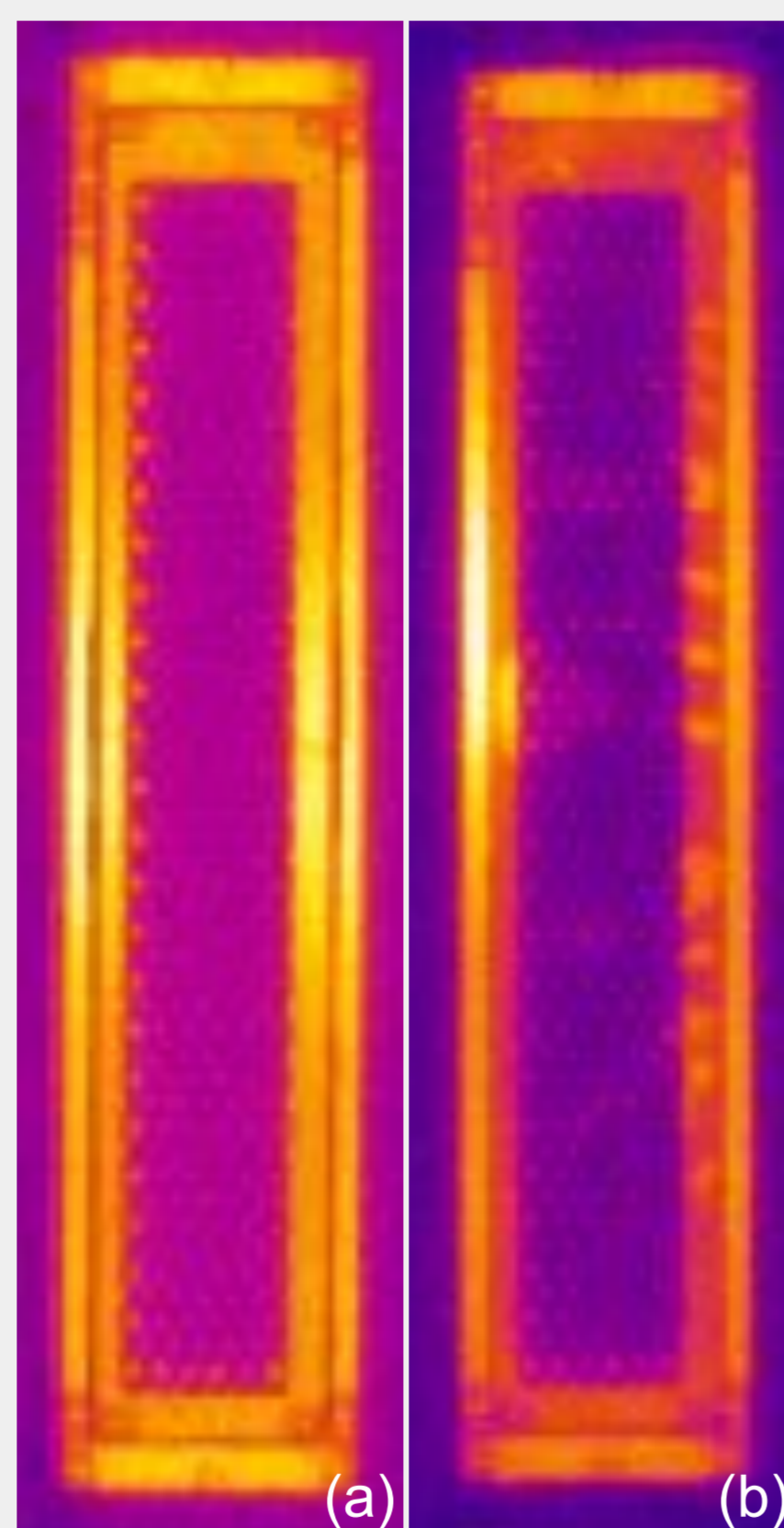
- Successful demonstration of hybridization
  - Microlens-detector array bonding
  - Precision alignment better than  $10 \mu\text{m}$
  - Low-loss sub-micron thick epoxy layer
  - No voids or dust between dies
  - Quarter-wavelength Parylene-C AR coating
- FLIR ( $\lambda \sim 10 \mu\text{m}$ ) and Zygo imaging used to verify:
  - Absence of inclusions between dies
  - Bond layer thickness uniformity
- Successful repeated cryogenic cycling tests
  - Robust bond and no ARC delamination



Above: Photograph of  $\lambda \sim 25 \mu\text{m}$  hybridized microlens array. Below: FLIR photographs showing prototype hybridized arrays with (right) and without (left) inclusions.

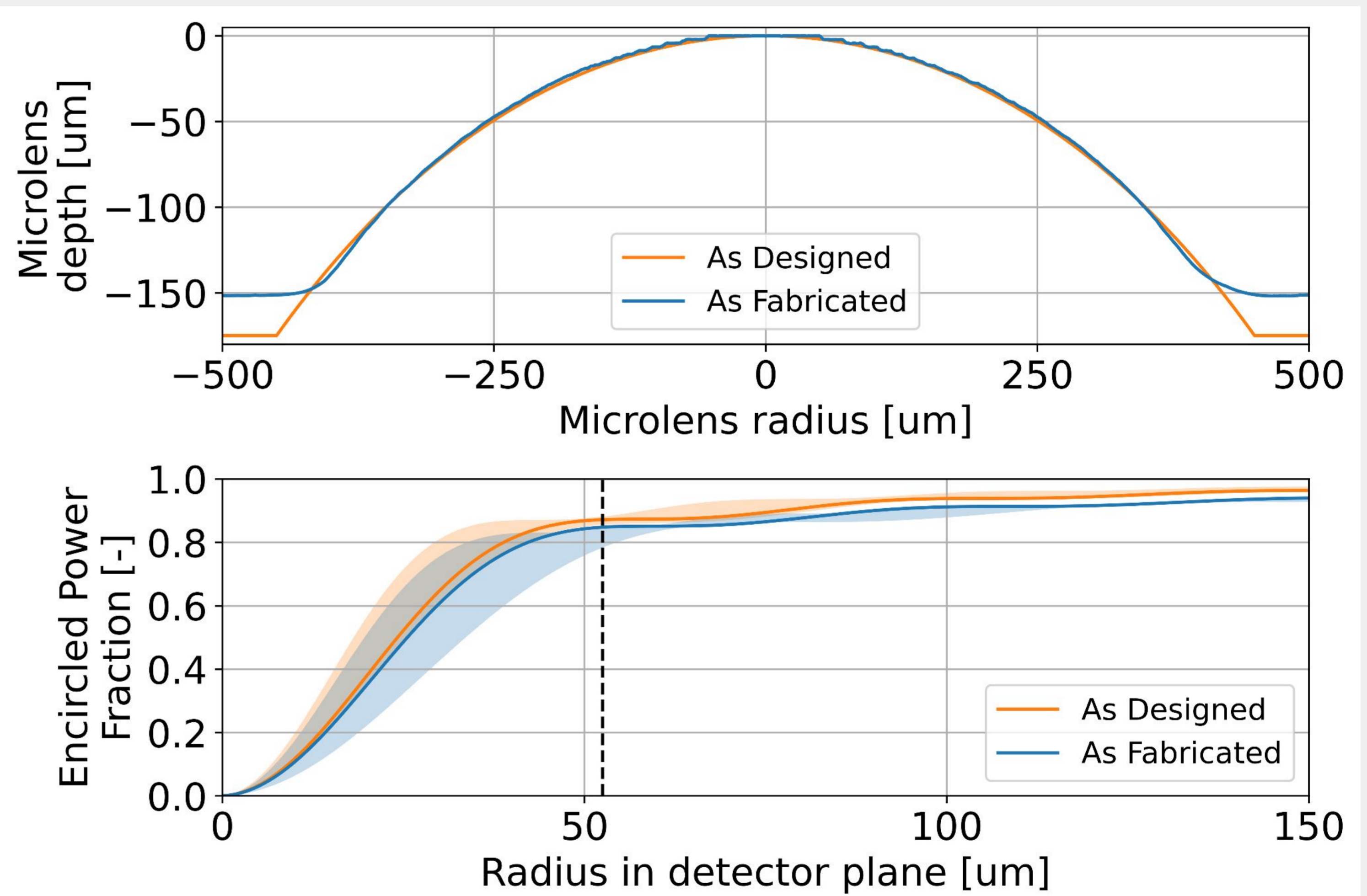


Above: Zygo optical profilometer measurement of hybridized array showing that the microlens die and detector die conform with a uniform bond thickness.



## Surface Accuracy and Estimated Efficiency of Fabricated Arrays

- Successful demonstration of lens figure and surface roughness
  - Microlens surface figure deviations  $\leq \pm 3\%$
  - Sub-micron surface roughness across lens
- Diffractive calculation to estimate expected encircled power
  - 85% efficiency for  $105 \mu\text{m}$  absorber diameter
- Microlens arrays currently in use for PRIMA detector development [1]



Upper plot: Designed versus fabricated profile of a prototype microlens designed for  $\lambda \sim 200 \mu\text{m}$  with a pitch of  $900 \mu\text{m}$ .

Lower plot: Calculated encircled power fraction for designed and fabricated microlens profiles. Solid lines indicate center wavelength ( $180 \mu\text{m}$ ). Shaded regions indicate performance spanning this PRIMA band ( $135 - 240 \mu\text{m}$ ).

## Conclusions and Next Steps

- Successful demonstration of monolithic microlens arrays
  - Fabricated prototypes exceed required performance
  - Hybridization method proven and robust
- Hybridized microlens arrays in use for PRIMA detector development [1,2,3]
  - $\lambda \sim 30 \mu\text{m}$  hybrids used for sensitivity measurements
- Next steps:
  - Fabricate and hybridize PRIMA brassboard demonstration arrays
  - Extend to use for other projects (BEGINS [4])

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